# Laboratory observation of forbidden x-ray transitions in solar wind charge exchange

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## 1. Introduction

In 1990, the ROSAT satellite was launched and measured the soft x-ray background over the all sky. During the observation, it was discovered that the intensity of the soft x-ray background varies with short time within a few days in space where x-ray sources do not exist [1,2]. In 1996, the ROSAT satellite observed the soft x-ray emissions from the Hyakutake comet when it approached the earth [3]. After that, the x-ray emissions from various comets were observed and it was concluded that a part of the emissions is caused by the de-excitations in charge exchange processes between highly charged ions (HCIs) in solar wind and cometary neutrals<sup>4)</sup>. Since then it has been recognized that the long-term enhancements of the soft x-ray background stem from the solar wind charge-exchange (SWCX) with neutrals within the heliosphere as well as cometary neutrals [5-7]. In order to quantitatively analyze the high-resolution x-ray spectra observed with the satellites in detail, the cross sections in SWCX processes, in particular x-ray emission cross sections, are needed. However, there are no sufficient cross section data so far.

In these contexts, some x-ray emission cross sections in the SWCX processes have been measured for the electric dipole allowed x-ray transitions [8,9]. However, the main emissions in SWCX are known to be the forbidden and inter-combination transitions in  $O^{6+}$  (1s2s/1s2p - $1s^2$  transitions) [7], which have not been measured yet in a laboratory. Since the ion beam has a velocity of 200 - 900 km/s in order to reproduce the solar wind, it is impossible to observe the transitions with long lifetimes over microseconds by the beam-based experiments. Nonetheless, we have a chance to observe the forbidden transitions by trapping the metastable HCIs after the charge-exchange reactions of  $O^{7+}(1s) + He \rightarrow O^{6+}$  $(1snl) + He^{+}$ .

In this work, we report on the development of a Kingdon ion-trap system for trapping high energy HCI beams produced by an electron-cyclotron resonance ion source (ECRIS). As a performance

test of the instrument, we measured trapping lifetimes of  $Ar^{q^+}$  (q = 5, 6) under a constant number density of H<sub>2</sub>. Some experimental problems and their solutions are discussed. Then, the charge-transfer cross sections of the  $Ar^{q^+}$  - H<sub>2</sub> collision systems were measured at binary collision energies of a few eV [10]. The results are compared to previous data and the values estimated by some scaling formula. Finally we report on the status of the observation of the forbidden x-ray transitions from the metastable O<sup>6+</sup> (1*snl*).



**Figure.1** The side view of the Kingdon ion trap from the ion beam axis.

#### 2. Experimental Setup

In this abstract, I describe the setup of trapping lifetime experiment. The highly charged ion beams were produced by ECRIS at an acceleration voltage of 6.0 or 8.0 kV. We selected the ions with a desired charge-state by an analyzing magnet. Then the ions were injected into a Kingdon ion trap for trapping experiments. The ion beam current was measured with a Faraday cup installed at the most downstream of the beam line before applying high voltage into the trap electrodes.

Figure.1 shows a schematic drawing of the Kingdon ion trap. It consists of a central wire electrode, a cylinder electrode and two end-cap electrodes. After pre-determined storage time, the trapped ions were ejected and a part of the ejected

ions are detected by a micro channel plate (MCP). The output pulse signals from the MCP are counted by a fast multichannel scaler. In the earlier study using the Kingdon ion trap, D. A. Church *et al.* reported that no product ions with lower charge state were stored after the charge-exchange collisions between  $Ar^{q+}$  and Ar [11]. However, our study showed that the ions produced by the charge-exchange collisions between  $Ar^{q+}$  and Ar [11]. However, our study showed that the ions produced by the charge-exchange collisions between  $Ar^{q+}$  and  $H_2$  were still trapped. Thus, we introduced the TOF measurement system in order to determine the trapping lifetime of HCIs. The data acquisitions were repeated  $10^3 - 10^4$  cycles depending on the ion beam current and storage time.

#### 3. Results and Discussion

Figure.2 shows the decay of trapped  $Ar^{q+}$  (q = 5, 6) ions as a function of the storage time in  $H_2$  gas at a pressure of  $1.24 \times 10^{-5}$  Pa. The decay rates of  $Ar^{5+}$  and  $Ar^{6+}$  were determined be 28(6) s<sup>-1</sup>, 67(6) s<sup>-1</sup>, respectively. In order to determine the charge-exchange cross sections from the decay rates, we need to know the average velocities of  $Ar^{q+}$  ions. For this purpose, we have performed extensive numerical simulations of ion trajectories trapped in the realistic Kingdon trap using the commercial software (SIMION8.0) and the average velocities of  $Ar^{5+}$  and  $Ar^{6+}$  were determined to be  $1.8(0.3) \times 10^4$  m/s and  $2.0(0.4) \times$  $10^4$  m/s, respectively. On the other hand, the number density of H<sub>2</sub> was determined by the ionization pressure gauge. Finally, the charge-exchange cross sections for the Ar<sup>5+</sup>-H<sub>2</sub> and Ar<sup>6+</sup>-H<sub>2</sub> systems were determined to be  $5.2(2.6) \times 10^{-15}$  cm<sup>-3</sup> and  $1.1(0.5) \times 10^{-14}$  cm<sup>-3</sup>, respectively. Note that the cross section of the Ar<sup>5+</sup>-H<sub>2</sub> system at the low energy was determined for the first time. These cross section data are consistent with the values estimated by the scaling formula and previous experimental data [12,13].

As a first step of the x-ray observation from the metastable  $O^{6+}$ , we started the trapping experiment. Figure.3 shows a decay curve of the trapped  $O^{6+}$  ions as a function of the storage time in H<sub>2</sub> gas at the pressure of  $4.4 \times 10^{-7}$  Pa. The trapping lifetime of  $O^{6+}$  is determined to be 280(15) ms. Since the lifetime of the forbidden transition in  $O^{6+}$  (1s2s - 1s<sup>2</sup> transition) is approximately 1 ms [14], the trapping lifetime is long enough to measure the forbidden transition.

The details of the discussions and the progress of the x-ray observation will be presented at the oral presentation and the master thesis.



**Figure.2** A plot of the extracted  $Ar^{5+}$  and  $Ar^{6+}$  as a function of storage time. The data are well fitted by single exponential function.



**Figure.3** A plot of the extracted  $O^{6+}$  as a function of storage time. This data is well fitted by single exponential function.

#### References

- [1] S. L. Snowden et al., Astrophys. J. 424 714 (1994).
- [2] S. L. Snowden et al., Astrophys. J. 454 643 (1995).
- [3] C. M. Lisse et al., Science 274 205 (1996).
- [4] T. E. Cravens, Astrophys. J. 532 L153 (2000).
- [5] T. E. Cravens, Science 296 1042 (2002).
- [6] V. A. Krasnopolsky et al., Space Sci. Rev. 113 271 (2004).
- [7] R. Fujimoto et al., Publ. Astron. Soc. Japan. 59 S133 (2007).
- [8] T. Kanda *et al.*, *Phys. Scr.* T **144** 014025 (2011).
- [9] H. Shimaya et al., Phys. Scr. T 156 014002 (2013).
- [10] N. Numadate et al., Rev. Sci. Instrum. 85 103119 (2014).
- [11] L. Yang et al., Phys. Rev. A 50 177 (1994).
- [12] M. Kimura et al., J. Phys. B 28 L643 (1995).
- [13] K. Okuno et al., AIP Conf. Proc. 360 876 (1995).
- [14] J. R. Crespo et al., Phys. Rev. A 58 238 (1998).